# Finite Element Model for Atmospheric IR-Absorption

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The updating concerns mainly the sensitivity analysis.

The model is not anymore monochromatic, the system matrix **K** is now calculated separately for the two contributions water vapor and **CO<sub>2</sub>**: **K=KH2O+KCO<sub>2</sub>**. The contribution of 0.04 % **CO<sub>2</sub>** is small as compared to the contribution of 3.5 % water vapor: for the global and annual mean evacuation the contributions of **CO<sub>2</sub>** and water vapor could be dealt with in a monochromatic model. However, for the sensitivity analysis the two contributions have to be dealt with separately.

The surface sensitivity for doubling  $CO_2$  concentration from 0.04% to 0.08% is now 0.03 °C.

Updated September 2014 . now including the MATLAB listing in appendix 5

#### Introduction

It has been argued by <u>Claes Johnson</u> that what is called "back-radiation of heat" by IPCC (International Panel for Climate Change) authors is not a physical phenomenon. [1, 2]

Heat propagates spontaneously only from warm to cold.

In this paper a finite element model is presented to deal with the one-way heat propagation by radiation in the atmosphere and thereby absorption and re-emission by IR-active molecules, in particular water vapor and CO<sub>2</sub>.

The main result of the paper is that the influence of CO<sub>2</sub> is negligible: doubling the present concentration of 0.04% to 0.08% ppm will give a temperature increase of 0.03 °C.

### One-way heat propagation according to Stefan-Boltzmann

In the **one-way** heat propagation formulation by radiation, a **pair** of surfaces is identified: surface i with absolute temperature  $T_i$  and surface j with absolute temperature  $T_j$ , and  $T_i > T_j$ .

The warmer surface  $\mathbf{i}$  emits a heat flux  $\mathbf{q}$  in W/m<sup>2</sup> towards the colder surface  $\mathbf{j}$ , which absorbs it:

For Ti >Tj: 
$$q(i \rightarrow j) = \sigma(Ti^4 - Tj^4)$$
 and  $q(j \rightarrow i) = 0$  (1a)

Stefan-Boltzmann constant :  $\sigma = 5.67 \text{ e-8} \text{ W/m}^2/\text{K}^4$ 

In the **two-way** heat propagation formulation by radiation each surface emits heat according to its own temperature and eventually absorbs heat according to the temperature of an other surface.

This was the hypothesis of Prevost from back in 1791.

Already at that time Fourier did not agree with the Prevost proposal. See <u>Matthias Kleespies</u> [3] for the history of back-radiation.

In terms of Stefan-Boltzmann, formulated a century later, the two-way heat propagation formulation between communicating surfaces is described as:

Surface i: emits 
$$\mathbf{q_i} = \sigma T_i^4$$
 and absorbs  $\mathbf{q_j} = \sigma T_j^4$   
Surface j: emits  $\mathbf{q_j} = \sigma T_j^4$  and absorbs  $\mathbf{q_i} = \sigma T_i^4$   
For  $T_i > T_j$ : from surface i towards surface j  
an algebraically **net** flux  $\mathbf{q_i} - \mathbf{q_j}$  (1b)

The two-way heat propagation formulation (1b) and the one-way heat propagation formulation (1a) seem to be identical. And indeed the same temperature distribution is obtained in practical applications, as has been shown in an <u>earlier paper</u> of the author [4].

But the two-way concept of Prevost gives rise to spurious absorption as shown in [4]. It should be avoided. IPCC authors use the wrong software based on the Prevost concept and claim huge absorption by the atmosphere.

In the case of **Ti** >**Tj** the surface **j** would emit a quantity of heat from cold to warm: a violation of the Second Law of Thermodynamics.

### Finite element model for the IR absorption in the atmosphere

We will derive in this section a set of simultaneous linear equations for heat radiation in a multilayer semi-transparent stack according to the one-way heat propagation formulation, using the finite element method (FEM).

It is an amelioration of an earlier <u>finite difference model</u> [4]. We model the semi-transparent atmosphere as a stack of N-2 grids consisting of fully transparent holes and fully absorbing wires with a normalized cross-section (1-fi) and fi, respectively. Such a grid can be considered as a very fine gauze.

The lower level of the stack at node 1 represents the surface of the planet, outer-space is represented by node N.

To each node **i** of the stack a temperatures  $T_i$  is allocated, converted into variables  $\theta_i = \sigma T_i ^4$ . The dimension of these variables is W/m<sup>2</sup> and indeed  $\theta$ , although variables related to temperature, sometimes represent heat fluxes. It follows from the context.

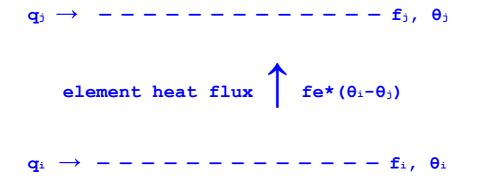
The wires of a grid in the stack exchange heat by radiation with the wires of other grids in the stack according to equation (1a), the one-way heat propagation formulation of the SB law, from the warmer ones to the colder ones, according to the Second Law of Thermodynamics.

In transverse direction *i.e.* parallel to the surface of the planet, the dimensions are supposed to be infinitely large: by symmetry only heat propagation by radiation in vertical direction has to be considered.

The dimensionless cross-section **f**i of the wire *i.e.* the diameter of the wire multiplied by its length per unit of area of the gauze, can vary for the various grids in vertical direction and can be interpreted as an absorption/emission coefficient. It will be called as such from now on.

We consider the one-way **heat** radiation between two layers **i** and **j** by LW radiation according to the SB law (1a), as depicted by the basic radiation finite element in **Figure 1**.

Figure 1 Basic finite element for heat radiation between grids i and j.



Grids i and j are not necessarily adjacent, other grids can be in between.

The nodal parameters are:

- absorption coefficients f
- to temperature related parameters  $\theta = \sigma T^4$
- thermal load into the element q

The constitutive relation of this radiation finite element is represented by the effective absorption:

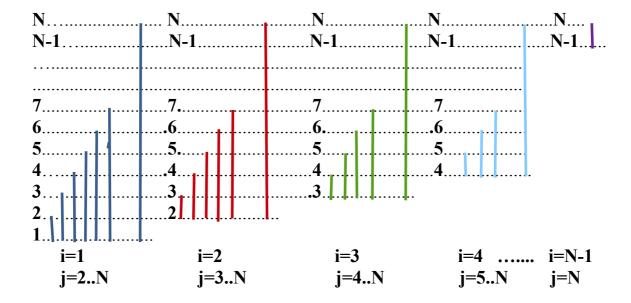
$$fe = f(i) *viewfactor(i,j) *f(j)$$
 (2)

The viewfactor of the element (i, j) expresses the window between two levels i and j, not necessarily adjacent. See **Appendix 2**.

By means of a Galerkin variation process (see **Appendix 1**) we find the basic radiation matrix for the finite element defined by nodes i and j in terms of the effective absorption fe of the element, giving a relation between the nodal parameters  $\theta$  representing temperature and the nodal thermal loads q:

For a stack of N levels, including the ground level 1 and the outer-space level N, there are N(N-1)/2 finite elements, which are assembled as shown in Figure 2, by superimposing elements of the same color ( with the same lower node) as well as elements of different colors (with different lower nodes), upon each other.

Figure 2 Illustrative color scheme of N(N-1)/2 finite elements (i , j) to be assembled by superimposing upon each other



The heat transport by LW radiation in the stack is therefore described by:

$$\mathbf{K}^*\boldsymbol{\theta} = \mathbf{rhs} \tag{4}$$

**K**: system matrix of order **NxN**, to be modified by boundary conditions.

The matrix **K** consists of two contributions: **K=KH2O+KCO2**.

The mean water vapor concentration is 3.5% and the mean **CO**<sub>2</sub> concentration is 0.04%. Because of the small concentration of **CO**<sub>2</sub> as compared to water vapor, for the analysis of the global and annual mean heat evacuation from the planet the difference in the distribution of water vapor and **CO**<sub>2</sub>, as given in figure **4**, has negligible effect. However for the sensibility analysis being the increase of surface temperature due to doubling the CO<sub>2</sub> contribution from 0.04% to 0.08%, the two contributions have to be dealt with separately.

 $\theta$ : vector of unknowns, W/m<sup>2</sup>, of order N

**rhs**: right hand side vector of fluxes **q**, W/m<sup>2</sup>, into the system, of order **N** 

In an earlier paper [4] the solution of (4) with boundary condition implemented by means of Lagrange multipliers is discussed. In the present paper we put the accent on the reversed solution technique.

### **Reversed Solution Technique**

When we know from observations the temperature distribution in the atmosphere then the reverse relation of (4) defines from known  $\theta$  and by means of the radiation matrix K, a vector  $\mathbf{q}$ :

$$q = K^*\theta$$
 (5)

In (5) the calculated vector  $\mathbf{q}$  represents the necessary heat input into the system to obtain the measured distribution of the temperature variables  $\mathbf{\theta}$ .

The calculated vector  $\mathbf{q}$  consists of the LW surface flux  $\mathbf{q}(1)$ , the term  $-\mathbf{q}(\mathbf{nods})$  is the **OLR**, outgoing LW radiation. The remaining components  $\mathbf{q}(i)$ , 1 < i < N, represent the heat deposited at the stations  $\mathbf{z}(i)$  by mechanisms other than LW radiation: absorption by aerosols of incoming SW, convection of sensible and latent heat, vertically by thermals and parallel to the surface by winds, etc.

These non-radiative mechanisms are necessary to compensate for the mismatch between emission and absorption for the measured temperature distribution defined by the environmental lapse rate ELR = dT/dz and the surface temperature Ts.

The measured environmental lapse rate **ELR** is between the dry adiabatic lapse rate **DALR** and the saturated adiabatic lapse rate **SALR**:

Dry adiabatic lapse rate: DALR = -g/cp = -9.81 K/km Environmental lapse rate: ELR = -6.5 K/km Saturated adiabatic lapse rate: SALR = -5 K/km

See Appendix 3.

We can write the components of the vector of temperatures T and of the corresponding vector of variables  $\theta$  as function of **ELR** and z:

$$T_i = T_S + ELR*z_i \rightarrow \theta_i = \sigma T_i^4$$
 (6)

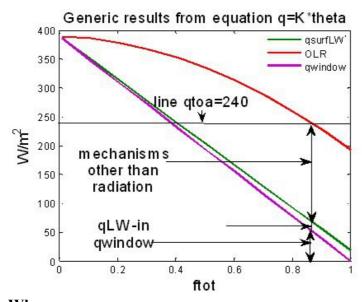
The vector  $\mathbf{q}$  in (5) represents now the necessary heat input at various height in the atmosphere to obtain a vector  $\mathbf{\theta}$  corresponding to a temperature distribution (6) according to the lapse rate **ELR**. We see that without a model for the various other mechanisms, the deposit of heat by such other mechanisms can be defined, not in detail of each mechanism but the detailed distribution in z-direction of the sum of all.

In view of the proverb "a picture tells more than thousand words", we give an example of the plots in the next sections where we will deal with the analysis of the atmosphere of the planet.

In **Figure 3** are plotted various heat fluxes and/or LW radiation as function of the parameter **ftot** which is the sum of the absorption coefficient of the grids. A typical value of the number of layers is **50**.

For a qtoa = 240 the value ftot= 0.86 is typical. For ftot =0 : OLR= qsurfLW=qwindow =  $\varepsilon 1*\sigma*Ts^4$ For increasing ftot, qsurfLW decreases: there is no back-radiation.

Figure 3 Results of equation  $q = K*\theta$ 



Where:

OLR: outgoing LW radiation = -q(nods) qsurfLW: LW surface flux = q(1) OLR - qsurfLW: mechanisms other than LW radiation = -q(nods)-q(1) qwindow: LW through atmospheric window = (1-ftot)  $*\epsilon(1)*\theta(1)$ 

**qLW-in** = **qsurfLW** – **qwindow:** LW flux into atmosphere

qtoa: outgoing LW flux at top of atmosphere, taken as 240 from IPCC.

ftot: sum(f)

Note the low value of the LW radiation into the atmosphere, **qLW-in**. Note also the difference between the red and green curve : *mechanisms other than LW* 

### **Input parameters**

### Geometry

The results as depicted in figure 3 are obtained from a FEM model with a height = 10 km and number of nodes N=50

The number of finite elements is N\*(N-1)/2=1225.

radiation. These mechanisms are predominant.

Such a high number allows to have a fine mesh near the surface where things happen. A fine mesh is needed there to be able to follow the strong variation of absorption near the surface.

The thickness of the layers dz(i) is defined by a geometrical series with ratio r=1.3 and with sum(dz) equal to height.

The first layer near the surface has a thickness of 6 mm, the upper layer has a thickness of 2.3 km.

### Distribution of absorption coefficients

The IR-active gases are those with three or more atoms per molecule: water vapor H20, carbon dioxide CO2, ozone O3, methane CH4.

Water vapor is the predominant IR-active gas: a concentration of **3.5%** as the annual and global average. The concentration of CO2 is 0.04%, being about **1%** of that of water vapor. The absorption by water vapor molecules is **10** times bigger, and as a consequence the contribution to absorption of CO2 is only **0.1%** of the total one defined by water vapor.

However, in the sensitivity analysis two contributions of CO<sub>2</sub> are analyzed: **0.1** % and **1**% of the total contribution of water vapor.

The distribution of the absorption of the water vapor over the various levels with thickness **dz(i)** is introduced by a function:

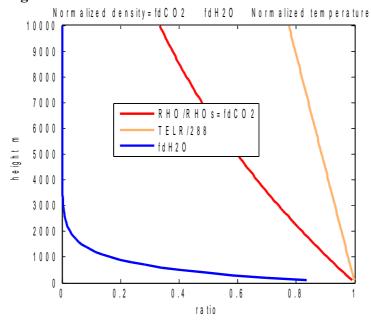
$$fdH_2O(i) = \alpha^*dz(i)^*exp(-m^*z(i)/(height/2))$$
(7)

where **m** is an input parameter.

The constant  $\alpha$  is such that the function is normalized: sum(fdH2O)=1.

The distributions fdH2O (7) and fdCO2 (8) are depicted in Figure 4.

Figure 4



For bigger **m** the IR-active water vapor is concentrated closer to the surface where the temperature is higher and as a consequence, the surface is emitting less LW for a given temperature according to the SB equation (1a).

It has turned out [4] that a value **m=9** give results corresponding to experimental results by IPCC such as the K&T diagrams.

In [4] the distribution of CO<sub>2</sub> was not specified, in the present FEM model we introduce the correct distribution of CO<sub>2</sub>, proportional to the density of air, since the volumetric concentration of CO<sub>2</sub> in the atmosphere is nearly constant.

In **Appendix 4** the density variation of air is defined as function of height **z(i)**, of the environmental lapse rate **ELR**, of gravitation constant **g** and of the gas constant **R**, from which follows a distribution function **fdCO2**:

$$fdCO2 (i) = \beta * dz(i)/(1+ELR*z(i)/Ts)^{(1+g/R/ELR)}$$
(8)

The constant  $\beta$  is such that the function is normalized: sum(fdCO<sub>2</sub>)=1. The distributions fdH<sub>2</sub>O (7) and fdCO<sub>2</sub> (8) are depicted in figure 4 The distributions of the absorption coefficients become:

$$fH20(i) = fdH2O(i)*(ftot-ftotCO2)$$

$$fCO2(i) = fdCO2 (i)*ftotCO2$$

$$f(i) = fH20(i) + fCO2(i)$$

$$(9)$$

NB ftot = sum(f), which follows from  $sum(fdH2O) = sum(fdCO_2)=1$ 

The parameter **ftot** is the total fraction of IR-active gases including contribution and distribution of water vapor, **ftotH2O**, and contribution and distribution of **CO2**, **ftotCO2**.

The absorption coefficients f(i) as defined in (9) are used in the constitutive relation (2) for the effective absorption fe of the various finite elements.

### Application to the real atmosphere.

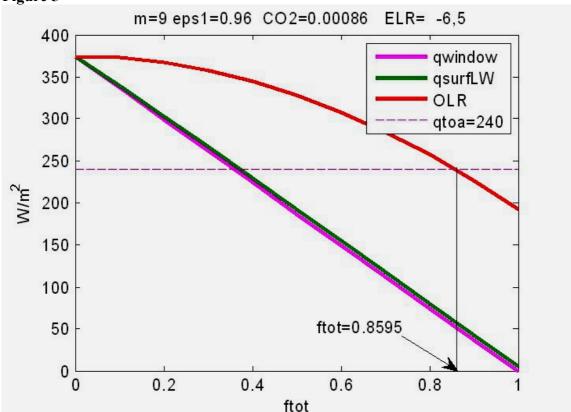
In **Figure 5** are given the results for runs with a mesh of **50** levels over a height of **10** km, in total **1225** elements, super imposed upon each other. The lower element has a thickness of 6 mm the highest element has a thickness of 2.3 km.

Water vapor distribution is according to (8) with m=9.

The results are given as function of **ftot**, mainly due to water vapor.

The absorption by CO2 is defined by fotCO2=0.00086 which is 0.1% of ftot.





For OLR=240 we find ftot = 0.8595.

LW radiation into the atmosphere is: qsurfLW - qwindow = 59-53=6

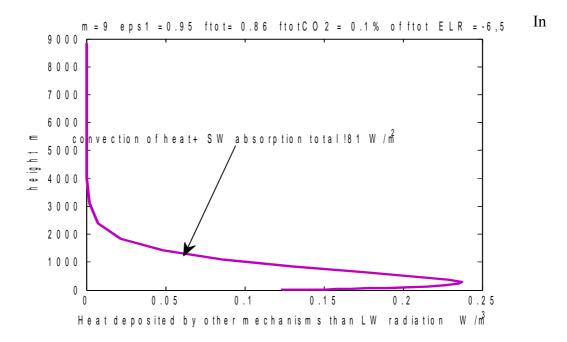
Mechanisms other than radiation produce: OLR - qsurfLW = 240-59 = 181 These mechanisms consist of:

109 as thermals of sensitive and (mostly) latent heat from the surface +72 as absorption of SW by aerosols, ice crystals etc in the atmosphere.

In **Figure 6** the distribution of **181**W/m<sup>2</sup> due to "mechanisms other than LW radiation" is plotted as W/m<sup>3</sup> by converting the nodal values of **q** taking into account the thickness

of the various finite elements. The surface flux q(1) = qsurfLW and the outgoing LW radiation OLR = q(nods) are not considered in this plot.

Figure 6

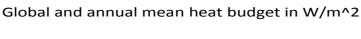


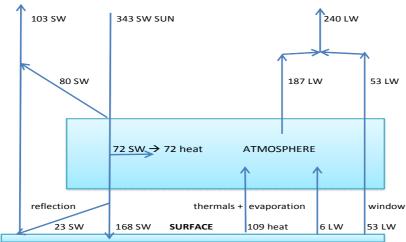
In **figure 7** is given a global and mean energy balance of the planet based on the present finite element model of the atmosphere.

Note the lack of back-radiation and the low value of  $\mathbf{qLW}$ -in =6 W/m<sup>2</sup> LW absorption by the atmosphere.

In an earlier paper [4] the value for qLW-in was 16. The difference is due to the viewfactor in the present FEM model.

Figure 7





### **Sensitivity Analysis**

As already been argued in a <u>foregoing paper</u> [4] with a finite difference implementation of the stack model, the temperature distribution in the atmosphere is defined by the environmental lapse rate **ELR** and the surface temperature **Ts.** 

Figure 5 for m=9,  $\varepsilon 1=0.96$ , ftot = 0.86, and qtoa = 240, characterizes the global and annual heat budget of the planet, depicted in Figure 7.

Parameter studies for other values of **m** have shown that the value of **m** has less effect on the pair (**qtoa**, **ftot**) but rather on the value for "mechanisms other than radiation".

We can conclude from this observation that the steady state temperature distribution will establish itself with the same global and annual mean lapse rate but a new global and annual mean surface temperature when the global and annual mean concentration of IR-active gases varies.

In Figure 5 we have shown results of **ftot** =0.86 with **ftotCO**2=0.00086 (= 0.1% of **ftot**) with distributions for water vapor and for CO2 according to (8) respectively (9).

From the sensitivity analysis we can also extract the results of the present stack model for the case that **ftotCO2=0.0086** (=1% of **ftot**).

We will give an analysis how the variation of the CO2 concentration will influence the surface temperature **Ts** according to the multilayer FEM stack model of the global and annual mean atmosphere.

From equation (8): 
$$q = K^*\theta = (KH20 + KCO2)^*\theta$$
 and  $OLR = -q(nods)$ 

With kH20nods and kCO2nods the last row nods of KH2O respectively KCO2:

$$qtoa = OLR = -(kH2Onods + kCO2nods) *\theta$$
 (10)

Variation of **qtoa** or **OLR** with respect to CO2 concentration and surface temperature depends on:

- variation of kCO2nods i.e the absorption value ftotCO2 and its distribution
- variation of  $\theta = \sigma T^4 = \sigma (Ts + ELR*z)^4$  i.e a variation of Ts assuming ELR is constant:  $\delta T_i = \delta (Ts + ELR*z_i) = \delta Ts$ .

We can write for a variation of **OLR**:

$$δOLR = -δ((kH20nods + kCO2nods)*θ)$$

$$= -δkCO2nods*θ - (kH2O nods + kCO2nods)*δθ$$
(11)

The sensibility of the surface temperature for doubling CO<sub>2</sub> concentration is defined by imposing  $\delta OLR = 0$ , assuming sun intensity and albedo do not vary for changing the CO<sub>2</sub> concentration from 0.04% to 0.08%. We can write:

$$B \Delta Ts + A \Delta ftotCO2 = 0$$
 (12)

The two parameters **B** and **A** follow from (11).

$$\theta_i = \sigma T_i^4 = (T_s + ELR * z_i)^4$$

$$\delta\theta_{i} = (4 \sigma T_{i}^{3}) \delta T_{i} = (4 \sigma T_{i}^{3}) \delta T_{s}$$
(13)

The parameter **B** becomes:

$$B = (kH2Onods + kCO2nods)*(4 \sigma T.^{3})$$
(14)

For Ts=288, ELR =-6.5 and ftot =0.86 the value becomes: B = 3.38.

The value for A depends only on the CO<sub>2</sub> concentration:

$$A = \Delta(kCO2nods*\theta) / \Delta ftotCO2$$
 (15)

 $(-Prevost + kco2nods*\theta)/\Delta ftotCO_2$ 

The term Prevost is the OLR for ftotCO2=0: Prevost =  $\varepsilon \sigma$ Ts0^4 = 374.48

For a value **fotCO2=0.00086**, **OLR** is calculated as:  $kCO2nods*\theta = 374.38$ 

The value A becomes: (-374.48 + 374.38)/0.00086 = -112.14

The constants A and B have been calculated for ftot=0.86 and ELR= -6.5 and Ts=288 K:

$$A/B = -33.216 \rightarrow \Delta T_S = -A/B*\Delta ftot CO_2 = 33.216*\Delta ftot CO_2$$

For 0.1% contribution of CO<sub>2</sub> or  $\triangle$ ftotCO<sub>2</sub>= 0.00086:

$$\Delta T_s = 33.216 * \Delta ftotCO_2 = 0.03 °C$$
 (16)

For 1% contribution of CO<sub>2</sub> or  $\Delta$ ftotCO<sub>2</sub> = 0.0086:

$$\Delta T_s = 33.216 * \Delta ftot CO_2 = 0.3 °C$$

#### **Conclusions**

The finite element model confirms the earlier conclusions of the <u>finite difference model</u> [4].

Once more it is shown that back-radiation does not exist nor the huge atmospheric absorption heralded by IPCC.

The sensitivity study for the surface temperature shows an increase of **0.03°C** for a doubling of the CO<sub>2</sub> concentration from 0.04% to 0.08%.

### Acknowledgments

The author wants to thank in particular <u>Claes Johnson</u> who inspired him to write this paper. The author interpreted his ideas by writing Stefan-Boltzmann always for a **pair** of surfaces: it opens the concept of standing waves.

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Thanks also to readers of previous versions of this paper for their suggestions how to ameliorate its message.

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- [2] http://claesjohnson.blogspot.fr/
- [3] http://principia-scientific.org/publications/History-of-Radiation.pdf
- [4] http://www.tech-know-group.com/papers/IR-absorption updated.pdf

### Appendix 1

### Concept of Galerkin variation to define element radiation matrix

Taking into account the fractional cross-section f(i) of grids, the one-way Stefan-Boltzmann law (1a) for heat flow between two grids with temperatures defined by  $\theta(i) = \sigma T(i)^4$  can be written as:

$$q(1\rightarrow 2) = f(1) *f(2) * (\theta(1) - \theta(2))$$
 (A1.1)

Figure A1.1 shows the basic radiation finite element with nodal parameters  $\theta$  and f related to the temperature respectively absorption/emission and with thermal loads q.

Figure A1.1 Finite element for radiation between adjacent grids

With fe = f(1) \* viewfactor(1,2) \* f(2) and viewfactor(1,2) = 1 the variational Galerkin statement for the first law becomes:

$$(\delta\theta(1) - \delta\theta(2)) *fe*(\theta(1) - \theta(2)) =$$

$$\delta\theta(1) *q(1) + \delta\theta(2) *q(2)$$
(A1.2)

The Galerkin variational statement is valid for arbitrary variations  $\delta\theta(1)$  and  $\delta\theta(2)$ , giving rise to two equations which in matrix form are written as:

### Appendix 2

### Viewfactor in an element with intermediate grids

In case that in an element, with end faces i and j, intermediate grids are present the exchange of heat by radiation is hindered by **viewfactor**(i, j). In case of one single intermediate grid: j=i+2:

The viewfactor reflects the fact that radiation from level i to level i+2 is hindered by the grid i+1.

In case more grids  $\mathbf{k}$  are present between level  $\mathbf{i}$  and level  $\mathbf{j}$ , with  $\mathbf{i} < \mathbf{k} < \mathbf{j}$ , the viewfactor of element  $(\mathbf{i}, \mathbf{j})$  becomes:

viewfactor(i, j) = 
$$1 - \sum_{k=i+1}^{k=j-1} f(k)$$
 (A2.1)

The so-called atmospheric windows from node i to outer space node N is a special case of the more general view-factors:

$$window(i) = viewfactor(i, N)$$
 (A2.2)

The effective absorption/emission of element (i, j) becomes:

$$fe(i,j) = f(i)*viewfactor(i,j)*f(j)$$
(A2.3)

### Appendix 3

### Lapse rates

The dry adiabatic lapse rate **DALR** is the thermal gradient **dT/dz** of dry air due the adiabatic compression/expansion of air.

It follows from the ideal gas law, being an equation of state for the 3 properties: pressure p [N/m^2], density  $\rho$  [kg/m^3], temperature T [K] and the gravitation law of Newton.

### Ideal gas law

$$p/\rho = RT$$
 with  $R = c_p - c_v$  and  $v=1/\rho \rightarrow pv = (c_p - c_v)T$ 

 $\begin{array}{lll} \textbf{R} & \text{specific gas constant} & [\text{Joule/kg/K}] \\ \textbf{c}_{\textbf{p}} & \text{specific heat for constant pressure} & [\text{Joule/kg/K}] \\ \textbf{c}_{\textbf{v}} & \text{specific heat for constant volume} & [\text{Joule/kg/K}] \\ \textbf{v} & \text{specific volume} & [\text{m}^3/\text{kg}] \\ \end{array}$ 

In differential form the ideal gas law is written as:

$$vdp + pdv = (c_p - c_v)dT$$

Two equivalent expressions result for the First Law of Thermodynamics for an ideal gas with independent variables **T** and **p**, respectively, **T** and **v**:

(1) 
$$dQ = c_p dT - v dp$$
 or (2)  $dQ = c_v dT + p dv$ 

$$dQ \text{ external heat supply} \qquad [Joule/kg]$$

The two expressions are equivalent: when subtracted from each other, the dQ term drops and the differential formulation of the ideal gas law results.

#### **Gravitation law of Newton**

$$dp/dz = -\rho g \rightarrow dp = -(g/v)dz$$

## Dry adiabatic lapse rate

With the gravitation law of Newton and with the adiabatic condition dQ = 0 the first expression of the First Law for an ideal gas yields:

$$dQ = c_p dT - v dp = c_p dT + g dz = 0$$

The dry adiabatic lapse rate becomes:

$$DALR = dT/dz = -g/c_p [K/m]$$

The saturated adiabatic lapse rate **SALR** depends also on humidity and the latent heat. See Wikipedia, lapse rate.

The measured global and annual mean environmental lapse rate ELR, between DALR and SALR, depends also on convection, conduction, absorption of incoming SW, etc.

### Appendix 4

### **Density variation in the atmosphere**

For given environmental lapse rate **ELR**, we can define the variation of air density with height.

From the ideal gas law we find

$$p = \rho RT \rightarrow dp/dz = RTd\rho/dz + \rho RdT/dz$$

Inserting the gravitation law of Newton:

$$RTd\rho/dz = -\rho RdT/dz - \rho g$$

With the lapse rate:

$$dT/dz = ELR$$
 and  $T = Ts + ELR*z$ 

One obtains:

$$d\rho/\rho = -(ELR*dz/(Ts+ELR*z))*(1 + g/ELR/R)$$

Integration gives:

$$ln(\rho) = -(1 + g/ELR/R)*ln(Ts+ELR*z) + constant$$

The integration **constant** follows from the boundary conditions for z=0:

$$\rho = \rho s$$
 and T=Ts  $\rightarrow$  constant = ln( $\rho s$ ) + (1 + g/ELR/R)\*ln(Ts)

The density variation is therefore:

$$\rho/\rho s = 1/(1 + ELR * z/Ts)^{(1 + g/ELR/R)}$$

Introducing the negative of dT/dz as lapse rate:  $\gamma = -dT/dz = -ELR$ 

$$\rho/\rho s = (1-\gamma * z/Ts)^{(g/\gamma/R-1)}$$

In **Figure 4** of the main text the distribution is depicted.

## Appendix 5

Since October 2014 the Matlab listing is in a separate paper:

http://www.tech-know-group.com/papers/Reynen-MATLAB-listing.pdf